

## Orbital Motion

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## Orbital Motion

- What kind of force causes objects to move along a circular path?  
**Centripetal Force!**
- To determine the orbital speed of a mass that is orbiting around a larger central mass, we must first recognize that the centripetal force keeping it in orbit is actually the gravitational force, so...

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## Orbital Motion

$$F = G \frac{Mm}{r^2}$$

- And...  $F_c = \frac{mv^2}{r}$
- So it follows that:  $G \frac{Mm}{r^2} = \frac{mv^2}{r}$
- The orbital speed of a satellite, therefore, can be determined by using:

$$G \frac{M}{r} = v^2$$

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## Orbital Motion

- The previous example shows us how the orbital speed of a satellite depends on its orbital radius.
- The farther out a satellite is from Earth's surface, the slower it can go and still remain in orbit
- **How far out (what orbital radius) must a satellite have to remain in geosynchronous orbit?**

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## Kepler's laws of Planetary Motion

- **1<sup>st</sup> Law:** Planets move in ellipses around the sun, the sun being at one of the foci of each ellipse.
- **2<sup>nd</sup> Law:** The line from the sun to the planet sweeps out equal areas in equal times.
- **3<sup>rd</sup> Law:** The time taken to complete one full revolution is proportional to the 3/2 power of the mean distance from the sun.
  - *Or: the orbital period squared is directly proportional to the cube of the mean distance from the sun.*

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## Orbital Period, related to Kepler's third law of planetary motion...

- Assume a circular orbit...the average orbital speed is equal to the distance traveled by the planet or satellite in one orbital period:

$$v = \frac{d}{t} \quad \text{Which Means...} \quad v = \frac{2\pi r}{T}$$

- Using the relationship we previously deduced:

$$G \frac{M}{r^2} = v^2$$

- We can relate Newton's law of gravitation to Kepler's 3<sup>rd</sup> law of planetary motion...

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### Newton's Law of Gravitation & Kepler's 3<sup>rd</sup> Law of Planetary Motion

- Known:  $v = \frac{2\pi r}{T}$        $G \frac{M}{r} = v^2$
- Substitute:
- Solve:
- And we get the value of the proportionality constant for Kepler's 3<sup>rd</sup> law:

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### Weightlessness

- The forces that act on an astronaut (or anything else) in orbit in a spacecraft, some distance  $r$  from the center of the earth are:
  - His weight (gravitational force from Earth)
  - The normal force ( $N$ ) from the "floor" of the spacecraft.
- What is the net force acting on the astronaut?

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- $F_{\text{net}} = \frac{GMm}{r^2} - N = \frac{mv^2}{r}$
- $\frac{GMm}{r^2} - \frac{mv^2}{r} = N$

$$N = (m/r)\{(GM/r) - (v^2)\}$$

When we looked at Orbital motion, we showed that  $(GM/r) = v^2$ ...so our equation above mathematically shows that the Normal force that the astronaut feels is equal to **0 N**...

Therefore: "Weightlessness" implies not that there is no gravity, but rather that there is no reaction force exerting the upward force that we sense as our weight.

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